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- However, not yet measured... detector technology has been missing.

(low-E recoils lose only 10-20% to ionization or scintillation)

 Cryogenic bolometers and other methods proposed, no successful implementation yet

> Cabrera, Krauss & Wilczek Phys. Rev. Lett. 55, 25–28 (1985) (prehistory of CDMS detectors)

qR<1

Iong wavelength
"sees" all nucleons
simultaneously

O

○ N²

(up to few tens of MeV)

initial and final states are indistinguishable (coherence possible)

recoil ~ few tens of eV for targets of interest

Fundamental physics:

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 1-ton detectors reach only > ~3 GWt reactor power
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 the list gets much wilder.





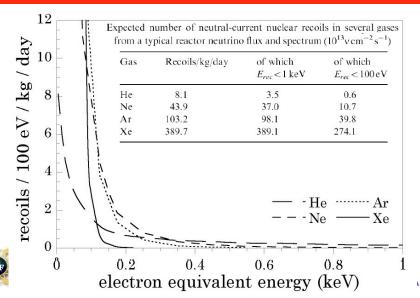
A one-page tutorial on coherent v-N scattering

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Detector mass must be at least ~1 kg (reactor experiment) + <u>recoil</u> energy threshold << 1keV

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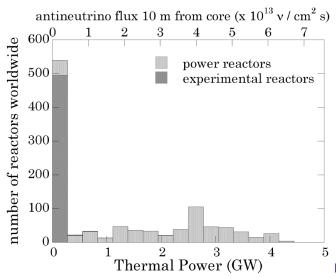




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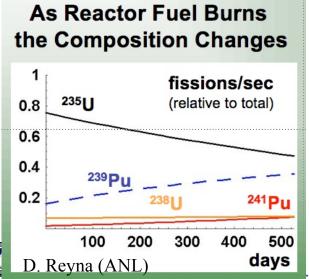




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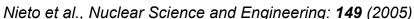


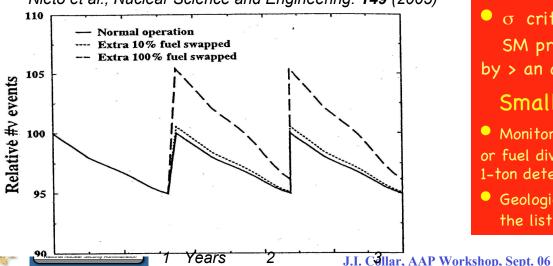


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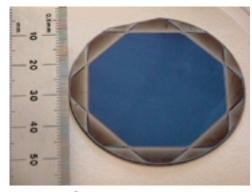
Kavli Institute for Cosmological Physics

Three legged stool needed: mass, threshold, background

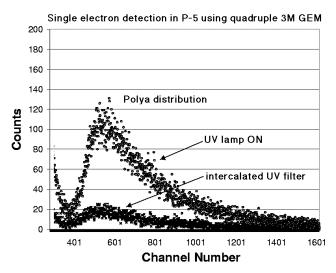
AT THE UNIVERSITY OF CHICAGO



No "light-bulb" moment: 5 years of R&D at UC



13 cm² APD on a 2 inch diameter silicon wafer



name-of-the-game:
detection of << 1 keV
recoils with large
(> 1 kg) detectors

(25 y and counting... must use new technologies or at least alterations)



Single-photon pulses using LN₂ cooled LAAPDs (high QE)

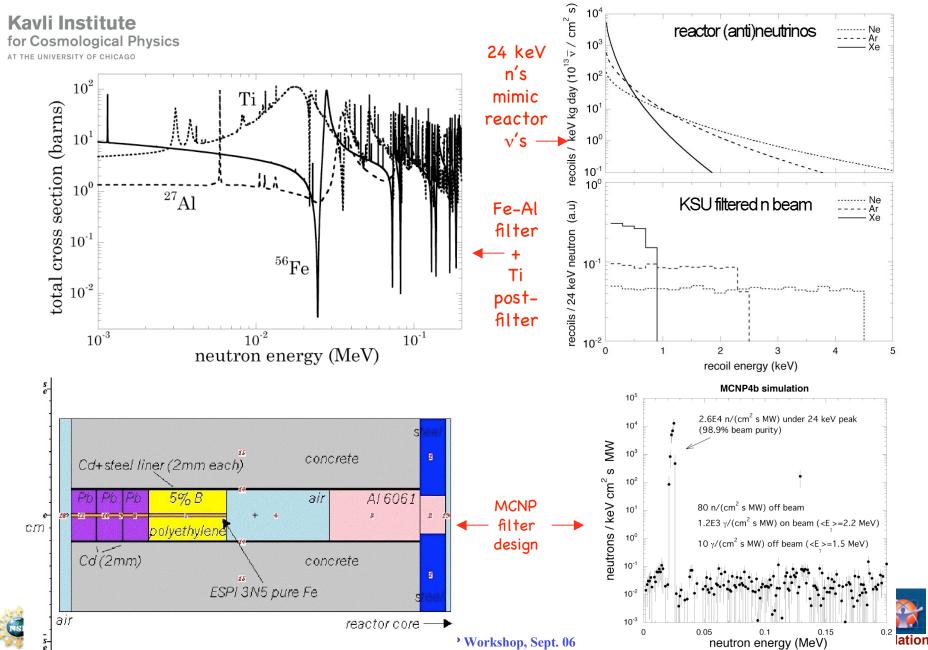








Start with the foundations: ultra low-energy recoil calibrations at KSU reactor





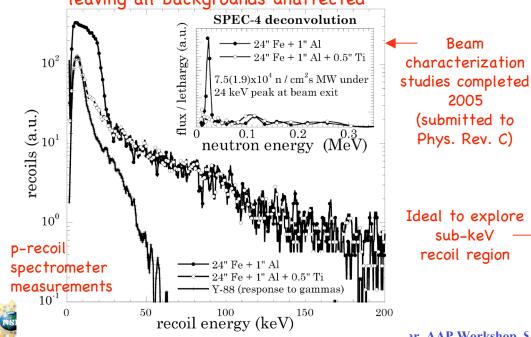
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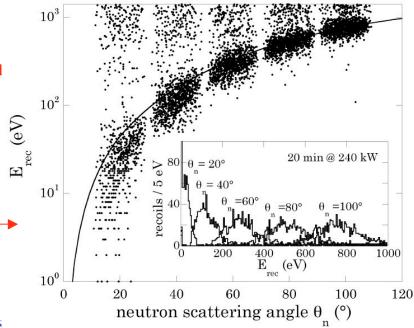








MCNP-POLIMI simulation 10^{3}



ar, AAP Workshop, §

Beam

2005

(submitted to

Phys. Rev. C)

sub-keV

recoil region



Modified-electrode p-type HPGe: A new tool in astroparticle & neutrino physics

Kavli Institute

IEEE Transactions on Nuclear Science, Vol. 36, No. 1, February 1989 P. N. Luke, F. S. Goulding, N. W. Madden and R. H. Pehl LOW CAPACITANCE LARGE VOLUME SHAPED-FIELD GERMANIUM DETECTOR 1989 state-of-the-art in large HPGe noise: 300 eV FWHM (even with modified electrode)

The idea: we have gone a long ways in JFET technology

2005: (factor x10 improvement in JFET C_F and V_n) ~50 eV FWHM (same as C_D ~1 pF x-ray detectors)

$$FWHM_{Ge} = 40.7 \ eV \cdot V_n (C_F + C_D) / \sqrt{\Delta t}$$

~1985 (TI 2N4416)
$$C_F = 4.2 \text{ pF}, V_n = 2 \text{ nV}/\sqrt{Hz}$$

2005 (EuriFET ER105)
$$C_F = 4.2 \text{ pF}, V_n = 2 \text{ if } V/\sqrt{Hz}$$

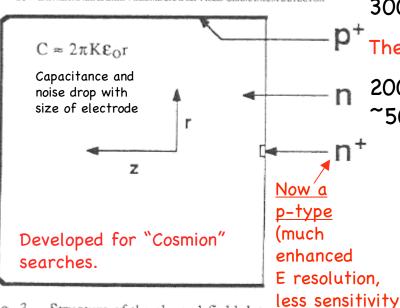


Fig. 3. Structure of the shaped-field detector.

backgrounds)









Modified-electrode p-type HPGe: A new tool in astroparticle & neutrino physics

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LOW CAPACITANCE LARGE VOLUME SHAPED-FIELD GERMANIUM DETECTOR $C \approx 2\pi K \epsilon_{O} r$ Capacitance and noise drop with size of electrode Z Now a p-type (much Developed for "Cosmion" enhanced searches. E resolution.

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~1985 (TI 2N4416)

 $C_F = 4.2 \text{ pF}, V_p = 2 \text{ nV}/\sqrt{Hz}$

2005 (EuriFET ER105) $C_{x} = 0.9 \text{ pF}, V_{y} = 1.6 \text{ nV}/\sqrt{Hz}$

Fig. 3. Structure of the shaped-field detector. less sensitivity backgrounds)

> The energy resolution and large mass of a HPGe plus the noise and threshold of a tiny x-ray detector???











 $C \approx 2\pi K \epsilon_{O} r$

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Capacitance and noise drop with size of electrode

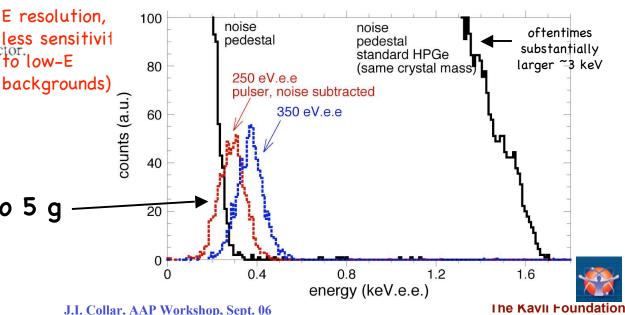
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~x10 less noise than conventional HPGe of same mass (475 g) (threshold equivalent to 5 g x-ray detector)

Fig. 3. Structure of the shaped-field detector.

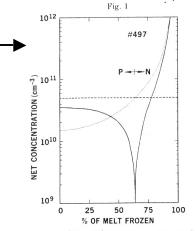








Precise gradient of charged impurities along the axis needed to compensate for small electrode



Net-shallow level concentration $|N_A - N_D|$ along the growth axis of an ultra-pure germanium single crystal. At the seed end (0% of melt frozen) the aluminium acceptor dominates, yielding the crystal p-type. Near the tail end the phosphorus concentration exceeds the aluminium concentration, |NA-ND|: continuous curve; aluminium concentration:

dashed curve; phosphorus concentration: dotted curve.

Delivered-To: collar@cfcp.uchicago.edu

Subject: Re: update?

From: otench@canberra.com

To: "Juan I. Collar" <collar@uchicago.edu>

Date: Sat. 3 Dec 2005 16:41:10 -0500

X-Uchicago-PMX-Id: 192.153.25.189; jB3LtheQ001154 [Sat Dec 3 15:55:44 2005]

X-Uchicago-Spam: Gauge=XXI, Probability=21%

Hello Juan,

We just got the first results in and they seem to be outstanding. The pulse resolution is about 160eV(FWHM) and Co-60 is well under 2.0 KV(FWHM). The detector should be shipping from France soon- in time for Christmas. It is too late now to change hardware but this might be done in future.

Best regards, Orren

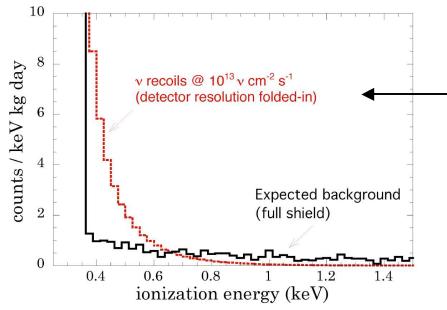
Developed during 2005 by CANBERRA/EURYSIS (the one of three contacted companies up to the challenge) Funded by NNSA.







Mass and threshold in place for reactor experiment, background... almost there (anti-Compton shield & Al part replacement underway)

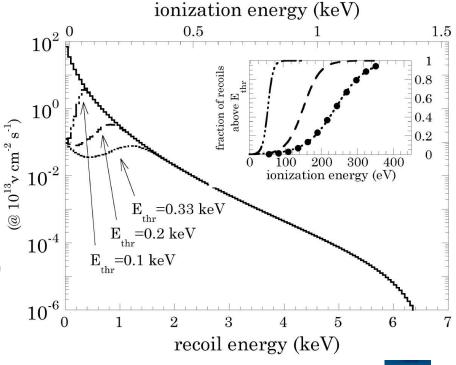


Expected antineutrino signal in reactor experiment with present detector. Background goal shown (scaled down from present status)

Presently 2.5 v recoils/kg-day expected

Work on non-white noise can increase
this to >30 v recoils/kg-day
(limited by state-of-the-art JFET noise only)

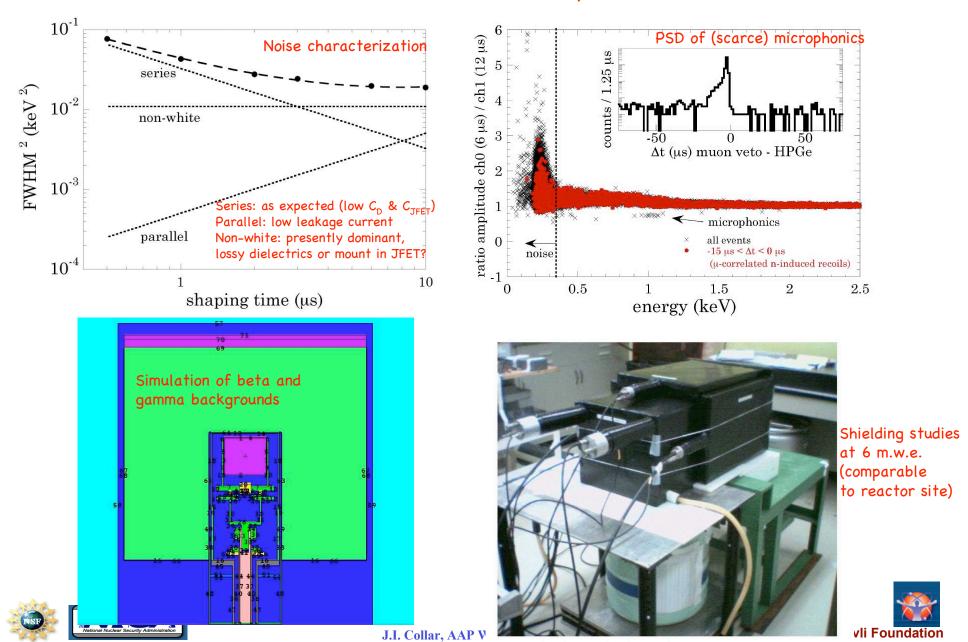
Silver lining: all of the signal concentrated in small ROI







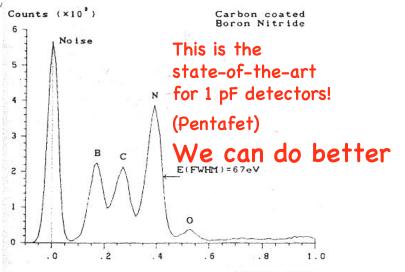
Extensive detector characterization early 2006 (submitted to Phys. Rev. C)

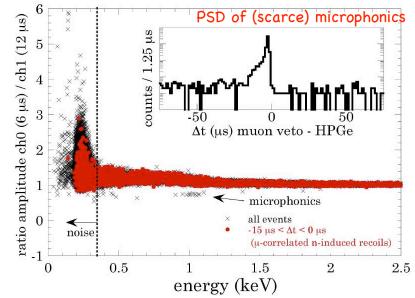


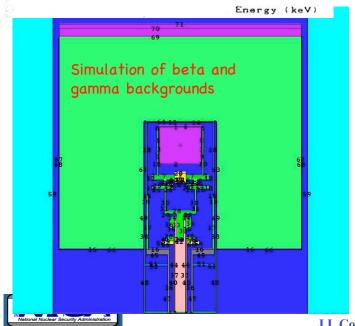


Extensive detector characterization early 2006 (submitted to Phys. Rev. C)

Kavli Institute for Cosmological Physics









Shielding studies at 6 m.w.e. (comparable to reactor site)

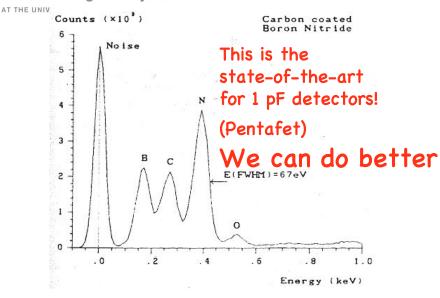


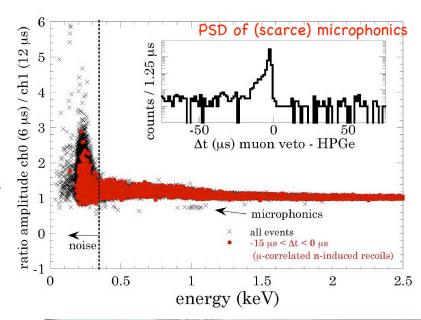


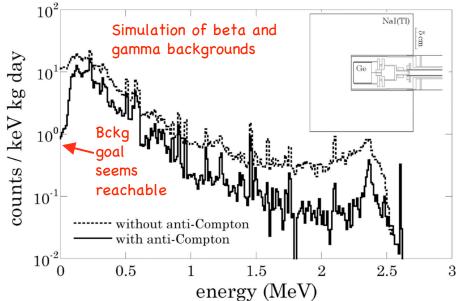


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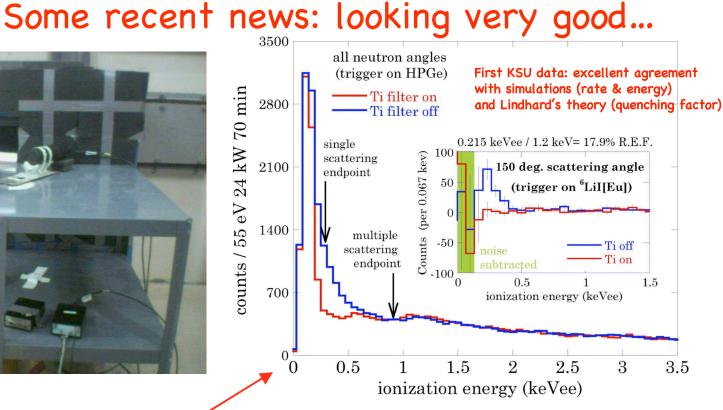


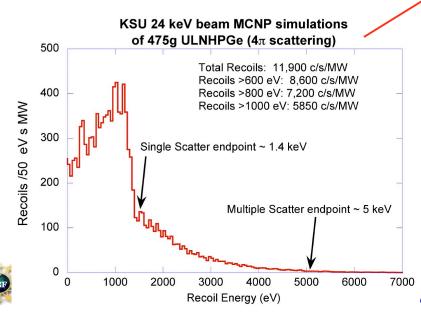


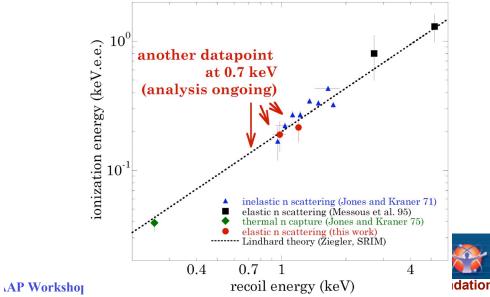








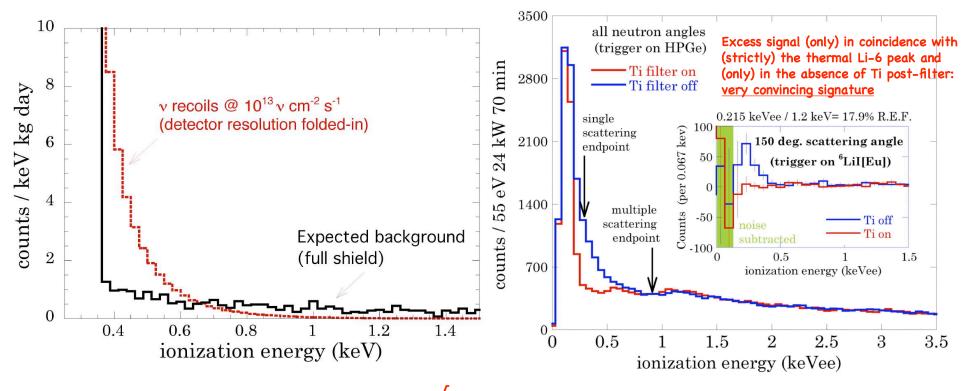






Some recent news: looking very good...

Relevant ROI for power reactor experiment has been explored



Mass $\sqrt{}$ Threshold $\sqrt{}$ Background... (on it)





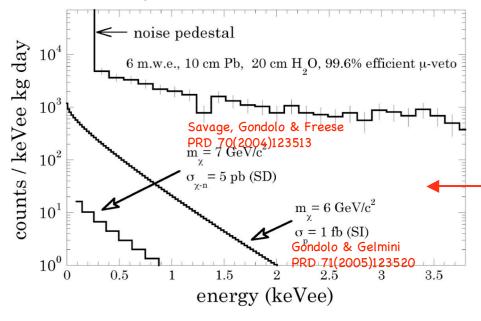




Some recent news: looking very good...

Kavli Institute for Cosmological Physics at the University of Chicago

Definitive check on DAMA soon X50 improvement from clean Al, x10 from anti-Compton



First Physics
Results expected
Fall 06:

These light WIMPs remain compatible with DAMA & all other searches (accelerator bounds are model-dependent)

Next: replacement with <0.2 ppb U cryostat, develop low-bckg version of anti-Compton shield... and deploy to power reactor.



Nothing radioclean yet (need cash! ⊗)







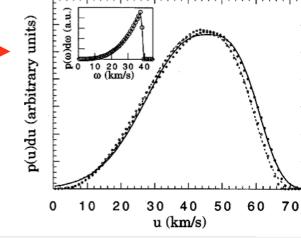


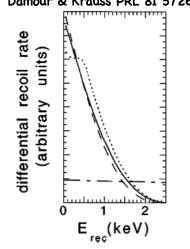
What else can you do with such a detector? J.I. Collar PRD 59 063514

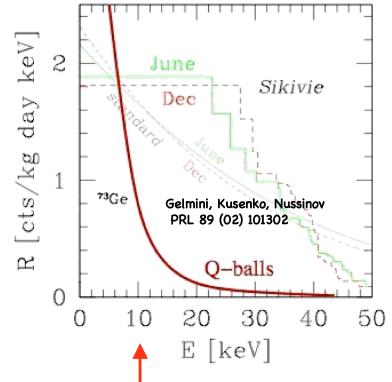
J.I. Collar PRD 59 063514 Damour & Krauss PRL 81 5726



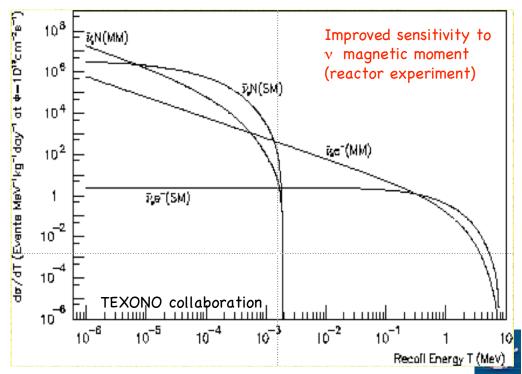
Solar-bound WIMPs:
deposit ~x100 less E_{rec}
than galaxy-bound,
concentrate all rate in narrow
spectral region (higher s/n)
Sub-keV threshold a must







The neutralino is not the only supersymmetric Dark Matter candidate. Non-pointlike DM (Q-balls, Mirror matter, etc.) call for ultra-low threshold detectors



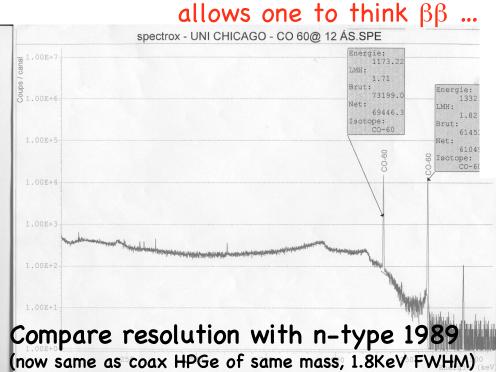


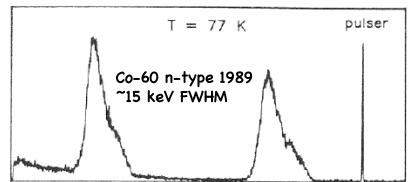




MAJORANA: can we avoid segmentation altogether? (cost, speed, simplicity, much lower front-end backgrounds) Does this device have anything to offer in a $\beta\beta$ context?

Optimal E-resolution





59.7

N+

P+

One of the position of the posit

Also x20 improvement in charge collection

in going to p-type

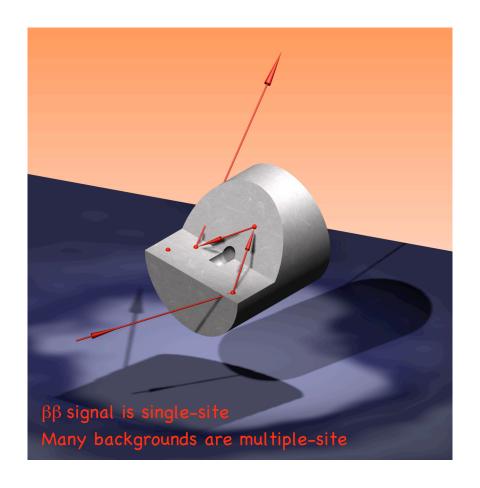








How does a multiple-site interaction look in a modified-electrode HPGe?



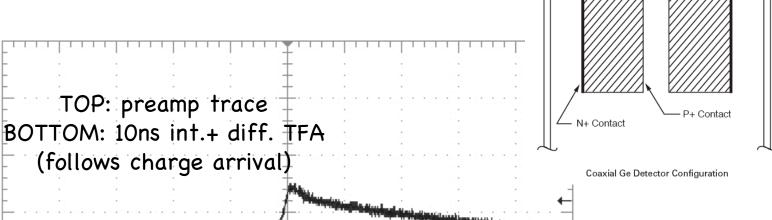


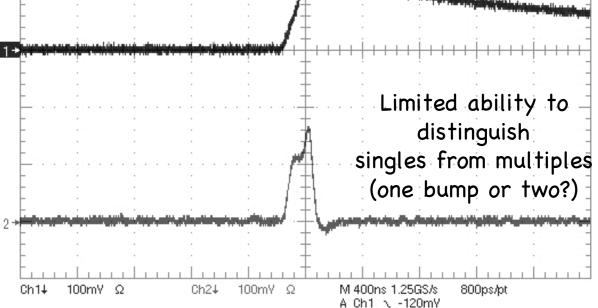






That was then...





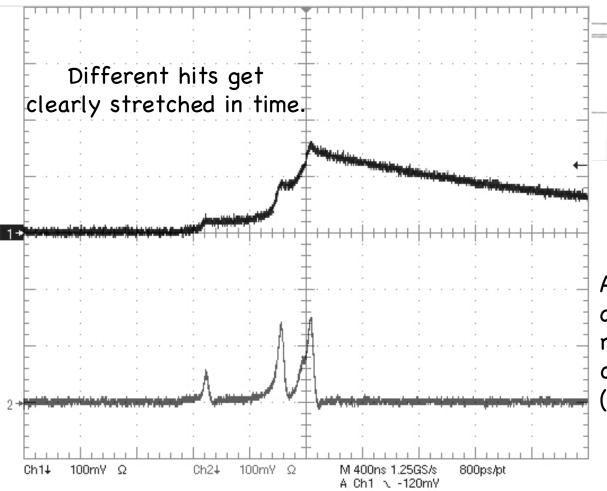


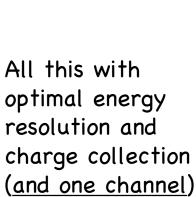






This is now.





Ø80

ø68



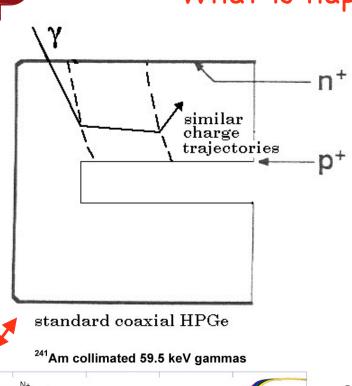


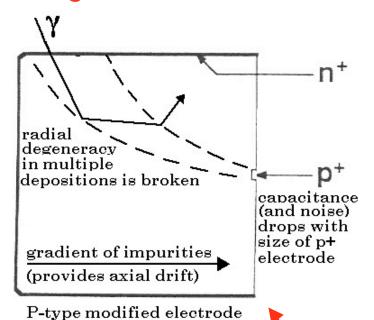


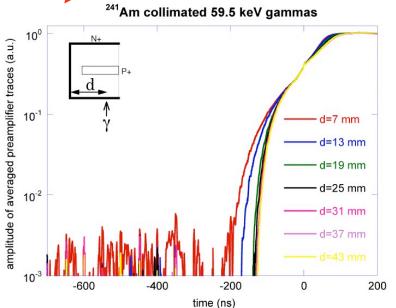
Kavli Instit

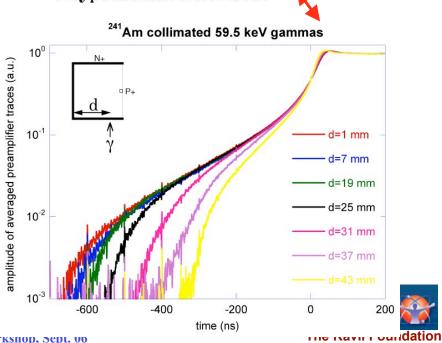
for Cosmologi
AT THE UNIVERSITY OF C

What is happening?







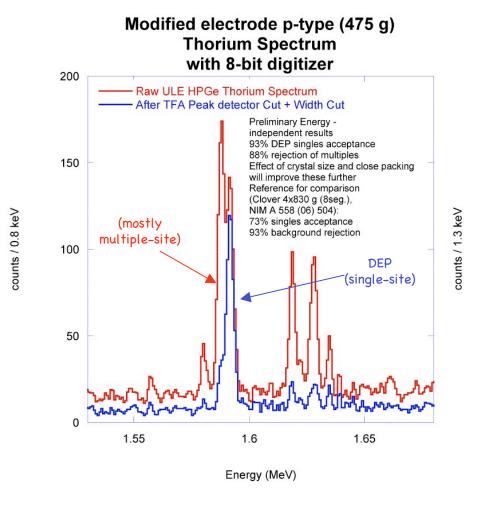


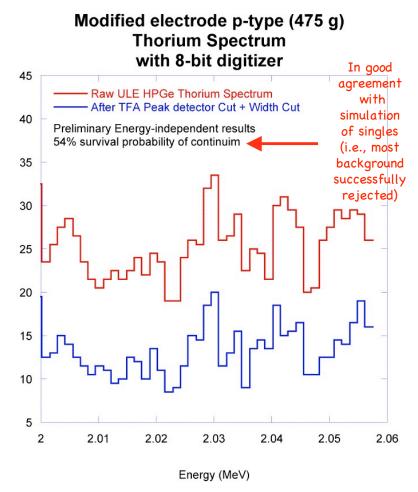






Better signal acceptance / background rejection than an 8-segment clover HPGe (even before close packing!) all with a single-channel device













Kavli Institute Advantages of single channel p-type modified for Cosmological Physics electrode vis-à-vis segmentation for MAJORANA:

- Very efficient PSA rejection of multiples. All with one channel.
- Excellent energy resolution (1.8 keV Co-60, may drop some as noise is further improved)
- Increase speed of deployment/manufacture <u>as long as</u>... (is gradient of impurities reproducible? How important?)
- Increase simplicity of construction and analysis (one channel)
- Decrease cost (detectors and DAQ). Improve production time (cosmogenics)
- Decrease front end-associated radioactive backgrounds, thermal load, photon path.
- Increase stability (prototype performance stable for >5 continuous mo. and counting)
- Intrinsic to p-type: ruggedness (a must when arraying) and decreased sensitivity to surface contaminations.
- Several others (e.g., rejection of ALL alphas via PIXE -studies underway-)
- CANBERRA and PHDs Co. receptive to further fabrication (and further work on noise reduction).







Disadvantages:

- Technology too new: too many unknowns in reproducibility, cost, speed of production, largest crystal size that can be produced, waste (important for $\beta\beta$), etc.
- Canberra's position: We need to build 6-10 more to know (they admit "lucking out".
 Hopefully this will not change).

Solution:

- This fits perfectly with planned coherent v program. Upcoming NSF proposal (3 yr) will be centered around this theme <- Help from rest of MAJORANA collaboration to maximize **synergy**: PNNL already funded to build more of these, ORNL seeking funding.
- Several kg of modified-electrode p-type HPGe's built by 2007!
- GOAL: Be by early next year counting at the Columbia Generating Station (Richland, WA, 12 mi. from PNNL) and simultaneously further developing the technique (i.e., building more of these). San Onofre? (offers more depth)







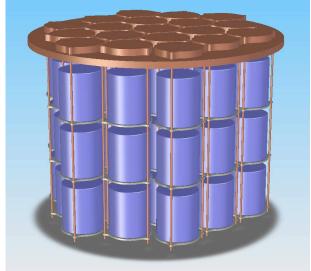


Reactor Monitoring: Right technological timing (HPGe technology flourishing:

Kavli Institute for Cosmological Segmentation, encapsulation, arrays and (silent) mechanical cooling)

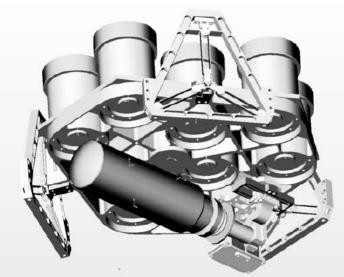
AT THE UNIVERSITY OF CHICAGO





MAJORANA

CLUSTER





GERDA







RHESSI

J.I. Collar, AAP Workshop, Sept. 06

The Kavli Foundation



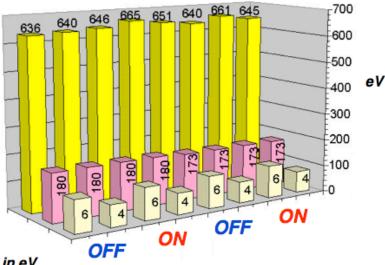
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Measurements with a BEGe2020 detector With CryoPulse 5 ON & OFF Vs shaping time



Resolution in eV

Temperature in negative ℃

Shaping time in µs

New generation of recondensing Dewars add no microphonic noise and need topping (not refilling) every ~ 1yr (can be filled from N2 gas cylinder!) Ideal for reactor deployment.





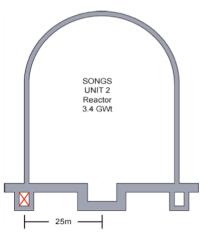




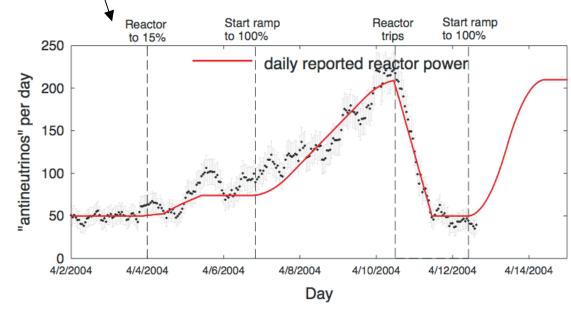
11 kg, encapsulated, single cold finger (CANBERRA)

A reality fast approaching?

With 100 eV threshold, the equivalent of ~1 ton liquid scintillator



SANDS Sees Reactor Turn-on in Detail (Antineutrino Rate, Running Average)





CLUSTER array for EUROBALL (7 encapsulated HPGe detectors)

Hexagonal tapering - diam.: 70 mm - height: 78 mm FWHM resolution : ≤ 2.3 keV

Efficiency: ≥ 55%

Alu wall thickness: 0.7 mm

Cap-to-Ge distance: 0.7 mm.







Coherent neutrino detection:



I want to believe!





